

Mark Scheme

Q1.

Question Number	Answer	Mark
	A	1

Q2.

Question Number	Answer	Mark
	B	1

Q3.

Question Number	Answer	Mark
	A	1

Q4.

Question Number	Answer	Mark
	D	1

Q5.

Question Number	Answer	Mark
	D	1

Q6.

Question Number	Acceptable answers	Additional guidance	Mark
	B		1

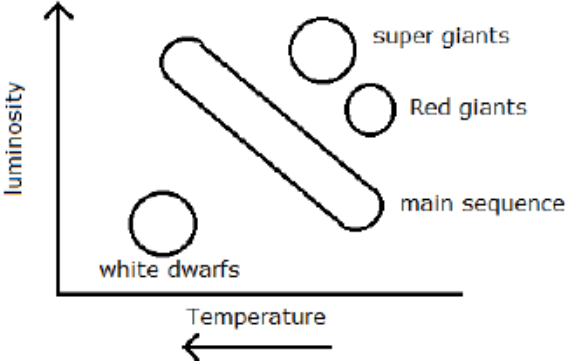
Q7.

Question Number	Answer	Mark
	B – (point on graph with luminosity $\neq L_{\odot}$)	1
	Incorrect Answers: A – luminosity $\neq L_{\odot}$ C – luminosity $\neq L_{\odot}$ D – luminosity $\neq L_{\odot}$	

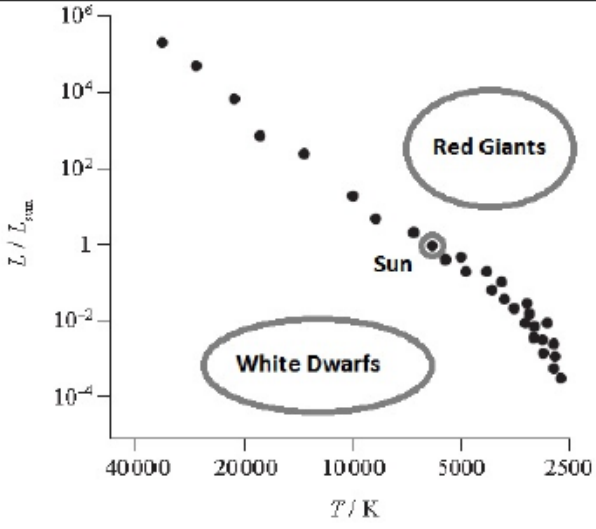
Q8.

Question Number	Acceptable answers	Additional guidance	Mark
	A description that makes reference to the following: <ul style="list-style-type: none"> • (main sequence) stars are (primarily) (1) converting hydrogen to helium in their core • stars on main sequence maintain a (1) constant luminosity (for most of their lifetime) 		(2)

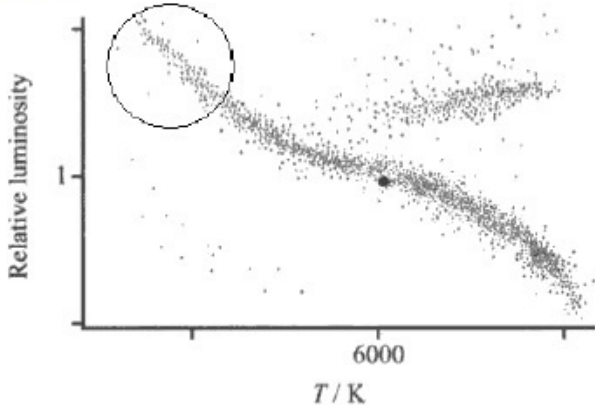
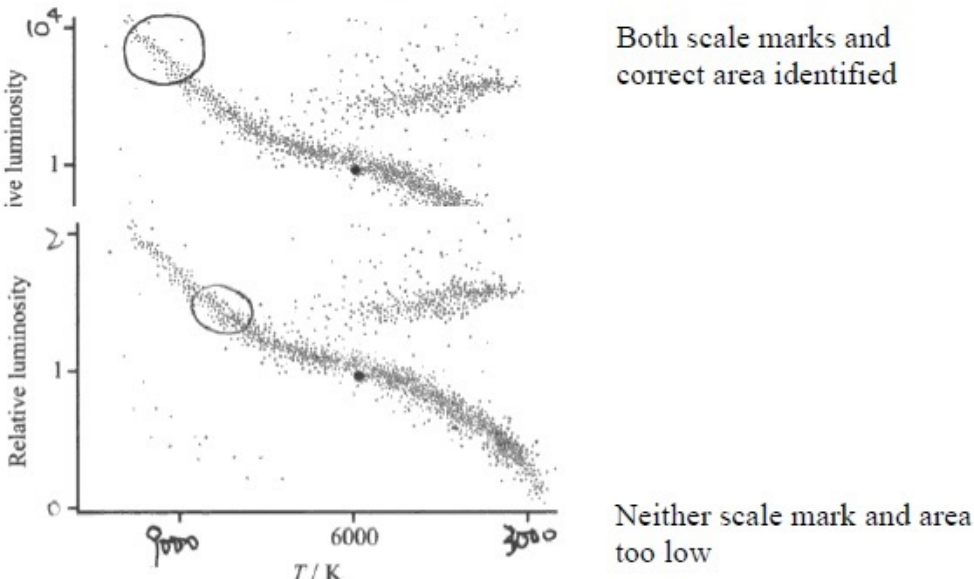
Q9.

Question Number	Acceptable answers	Additional guidance	Mark
	<ul style="list-style-type: none"> y axis: luminosity (/ luminosity of Sun) (1) x axis: (surface) temperature, with indication of decreasing temperature (1) 2 or 3 correct regions (1) 4 correct regions (1) 	<p><u>Example of graph:</u></p> 	(4)

Q10.

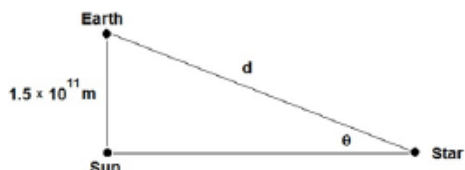
Question Number	Answer	Mark
		
(i)	Sun's position identified [single point identified]	(1)
(ii)	White dwarf region	(1)
	Red giant region	(1)
		3
*(iii)	<p>(QWC – Work must be clear and organised in a logical manner using technical wording where appropriate)</p> <p>White dwarf stars have:</p> <ul style="list-style-type: none"> high temperature T (because λ_{\max} is small) (1) low luminosity L (1) $L = \sigma AT^4$ linked to a determination of the surface area (1) 	3

Q11.

Question Number	Answer	Mark
(a)	Luminosity scale: Log scale [$10^3 \rightarrow 10^6$ (top) and $10^{-3} \rightarrow 10^{-6}$ (bottom)] (1) Temperature scale: reverse log/power scale [e.g. 12,000 (left) and 3000 (right)] (1)	2
(b)(i)	(Fusion of) hydrogen into helium [accept symbols] (1)	1
(b)(ii)	Circle around stars top left of main sequence [included in the area indicated below] (1)  Max 2 They have the highest temperatures Or they are the most luminous [accept brightest] (1) (Because) they fuse H (into He) at the highest/higher rate (1) (Because) they have the largest/larger gravitational forces (1) [Max 1 mark if no comparative] 	3

	Total for question	6

Q12.

Question Number	Answer	Mark
(a)	<p>Max 6</p> <p>The young star cluster consists (mainly) of main sequence stars (1)</p> <p>The old star cluster has a truncated main sequence (1)</p> <p>The old star cluster has lost its heaviest main sequence stars (1)</p> <p>The old star cluster has (many) red giant stars (1)</p> <p>The old star cluster has (some) white dwarf stars (1)</p> <p>Massive main sequence stars are the first stars (to deplete sufficient hydrogen in their core) to evolve into red giant stars. (1)</p> <p>Some red giant stars have evolved into white dwarf stars in the old cluster (1)</p>	6
(b)(i)	Star A is closer to Earth than Star B (1)	1
(b)(ii)	 <p>Use of appropriate trigonometric relationship (1)</p> $d = 4.0 \times 10^{16} \text{ m}$ <p><u>Example of calculation:</u></p> $\sin \theta = \frac{1.5 \times 10^{11} \text{ m}}{d}$ $d = 4.01 \times 10^{16} \text{ m}$	2
(c)	$\lambda_{\max} = 1.0 \times 10^{-6} \text{ m}$ (1) <p>Use of $\lambda_{\max} T = 2.9 \times 10^{-3}$ (1)</p> $T = 2900 \text{ K}$ (1) <p><u>Example of calculation:</u></p> $T = 2.9 \times 10^{-3} \text{ m K} / 1.0 \times 10^{-6} \text{ m} = 2900 \text{ K}$	3
	Total for question	12

Q13.

Question Number	Answer	Mark
(a)(i)	Reverse direction for temperature [at least 2 values seen] (1) Logarithmic/power temperature variation [at least 3 realistic values seen increasing by the same factor] (1)	2
(a)(ii)	QWC – Work must be clear and organised in a logical manner using technical wording where appropriate Area 1: Max 2 The Sun is fusing/burning hydrogen (into helium in its core) (1) When (hydrogen) fusion/burning ceases the core of the Sun cools [accept radiation pressure drops when fusion/burning ceases in the core] (1) (1) The core collapses/contracts (under gravitational forces) Area 2: Max 2 (1) The Sun expands and becomes a red giant (1) The core becomes hot enough for helium fusion/burning to begin (in the core) Helium begins to run out and the core collapses again (under gravitational forces) (1) Area 3: Max 2 (1) Idea that outer layers of Sun are ejected into space (1) The temperature doesn't rise enough for further fusion to begin (1) The core/Sun becomes a (white) dwarf star	6
(b)(i)	Idea of a very high temperature [accept value of about 10^7 K] (1) To overcome repulsive/electrostatic forces between protons/nuclei Or so that protons/nuclei get close enough together for the strong (nuclear) force to act Or so that protons/nuclei get close enough to fuse (1) Idea of a very high density [accept pressure] to give a sufficient collision rate (1)	3
(b)(ii)	Attempt at calculation of mass deficit (1) (1) Use of $\Delta E = c^2 \Delta m$ (1) Attempt at conversion from J to (M)eV (1) $\Delta E = 12.9$ (MeV) [If correct mass defect in kg is converted into u and then $1u = 931$ Mev used, then full marks may be awarded] <u>Example of calculation</u> $\Delta m = ((5.008238 \times 2) - 6.646483 - (1.673534 \times 2)) \times 10^{-27}$ kg	4

$\Delta m = 2.2925 \times 10^{-29} \text{ kg}$ $\Delta E = (3.00 \times 10^8 \text{ ms}^{-1})^2 \times 2.2925 \times 10^{-29} \text{ kg} = 2.063 \times 10^{-12} \text{ J}$ $\Delta E = \frac{2.063 \times 10^{-12} \text{ J}}{1.60 \times 10^{-13} \text{ JMeV}^{-1}} = 12.9 \text{ MeV}$	
Total for question	15